

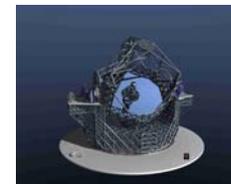
# TEOPS: "Groups, Commonality, Activity, Capabilities & Silicate bonding"

- What is TEOPS?
- Technology for Experimental and Observational Physics in Scotland
- Initiative in SUPA Astrophysics and Space Research theme
- "Spans the areas of particle physics, astrophysics and astronomy with a common theme of leading edge technology"
- Collaboration between UK ATC and Glasgow University Institute for Gravitational Research and Experimental Particle Physics groups

# Institute for Gravitational Research (IGR)

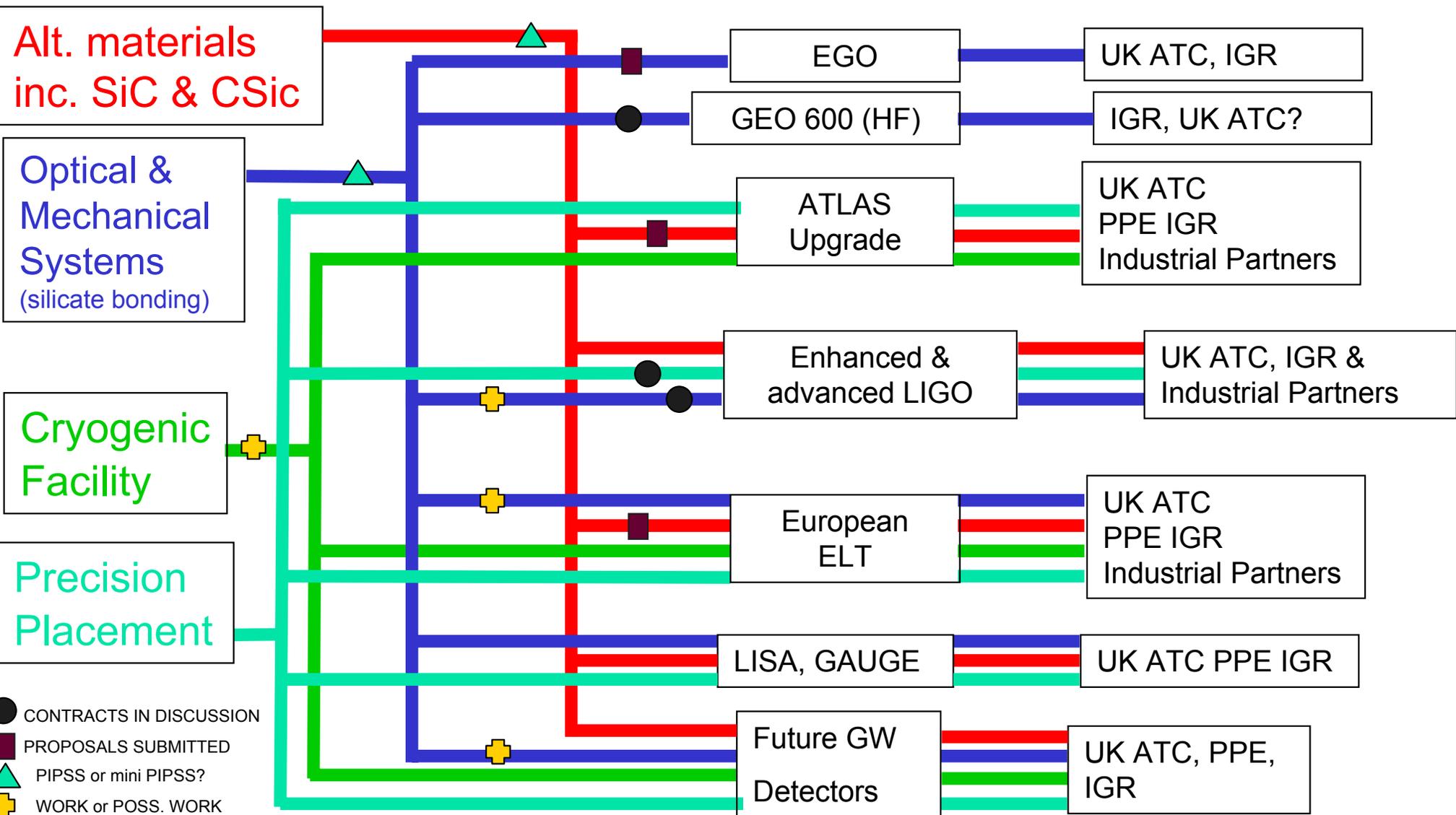
- *The work of the Institute is focused on*
  - *the development of detectors to search for gravitational waves from astrophysical sources*
  - *data analysis activities within the LIGO Scientific Collaboration.*
- *The main areas of experimental research are*
  - *development of precision novel interferometric techniques*
  - *development of systems of ultra low mechanical loss for the suspensions of mirror test masses along with research towards the space-based LISA mission*
    - Development of multiple pendulum systems using silica fibres to support the test masses
    - New bonding technology (hydroxide-catalysis bonding), which exhibits very low mechanical loss and is compatible with ultra-high vacuum
- *The technology developments within the IGR are of broader relevance to a number of areas of current PPARC interest and extensions of the bonding technology are being pursued with general application to precision optical systems on the ground and in space.*

- Experimental Particle Physics - Detector Development group
  - involved in a wide range of projects related to imaging, radiation detection and detector development, within particle physics, medicine, biology
- *Examples of current/recent projects include:*
  - CERN ATLAS - Production and testing of modules for the LHC/ATLAS
  - CERN Medipix - High sensitivity X-ray imaging for medical and synchrotron applications
  - Retinal imaging - Measuring the electrical activity of retinal tissues



- UK Astronomy Technology Centre, Edinburgh
  - MIRI - hosting the European PI and opto-mechanical design leads for this key instrument on the JWST, successor to the Hubble Space Telescope
  - Involved in building instrumentation for the most exciting international ground-based and space-borne astronomy projects
    - WFCAM - the largest infrared camera ever built, a cryogenic instrument now undertaking unique surveys in the Northern Hemisphere skies
    - SCUBA2 - the successor to SCUBA, one of the most successful ground-based instruments ever built, utilising a new generation of sub-millimetre CCD-like detectors
    - European Extremely large telescope (E-ELT) - The UK ATC is leading the UK's work towards an optical and infrared telescope of up to 42 m in diameter, recently approved into the design phase by ESO Council
- ATC is also involved in several UK and European network and technology development initiatives
- Technology and research strengths complementary to and highly relevant to IGR and PPE groups in Glasgow

# AREAS OF COMMONALITY/ INTEREST



- VMC (HAAS)

- 5 axis



- LISA Pathfinder lab

- precision assembly



- SRDG - Optical Characterisation Suite

- Detector Characterisation Suite (PPE)

# Optical Characterisation Suite

- Zygo interferometer
  - Suitable for surface flatness and curvature measurements on mirrors
  - Flatness measurements of  $<1$  nm are possible over a 4 or 6 inch wafer
- Wyko NT1100 optical surface profiler
  - Field of view from  $3.5 \times 4$  mm to  $50 \times 50$   $\mu\text{m}$
  - Step heights up to several mm
  - Surface roughness measurements down to 0.5 nm
- TM1000 tabletop SEM
  - $\sim 30$  nm resolution on insulating substrates as well as conductive (reduced charging)
  - Totally self contained no additional pumps
  - $\sim 2$  min pump down



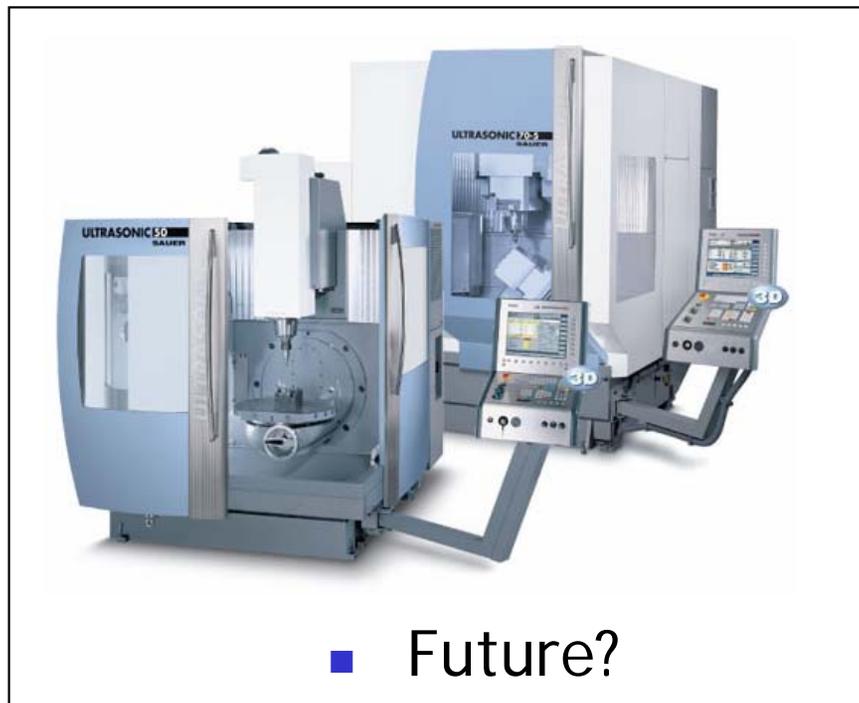
# Detector Development Lab (PPE)

- Automatic wire-bonding system
  - used for high precision, high density connections between detectors and readout electronics
  - used for the wire-bonding of SCT modules for the ATLAS project at CERN (~1.5 Mbonds)
- Cascade Microtech S300 probe station
  - sets the measurement standard for 300mm on-wafer test
  - applications
    - device characterization and modeling
    - wafer-level reliability
    - design de-bug
    - IC failure analysis
  - has the precision and versatility needed for the most advanced semiconductor processes and aggressively scaled devices
  - allows the group to create the perfect on-wafer measurement environment



# TEOPS: Capabilities #4

- Cryogenic Material Property Test Bed
  - Based at UK ATC

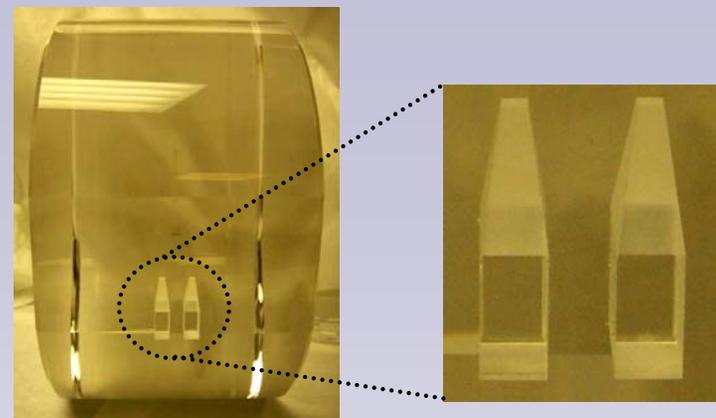




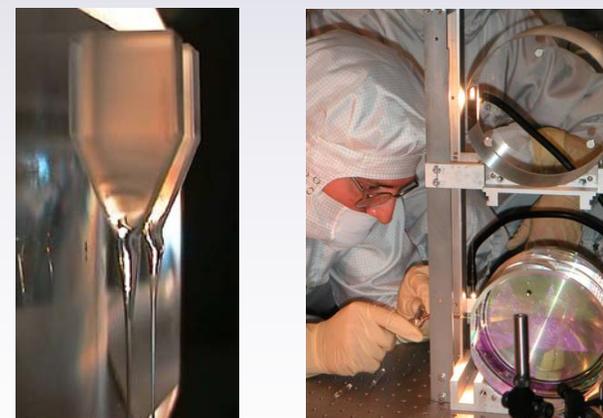
# Investigations into hydroxy-catalysis bonding

## Current applications

- Originally developed for NASA's Gravity Probe B mission, launched April 2004. (Gwo et al.)
- GEO600 currently operates with quasi-monolithic fused silica suspensions and mirrors. This **technology allows improved thermal noise** in the suspension systems.
- Construction of the ultra-rigid, ultra-stable optical benches for the LISA Pathfinder mission.



Picture of a GEO600 sized silica test mass in Glasgow with silica ears jointed using hydroxy-catalysis bonding



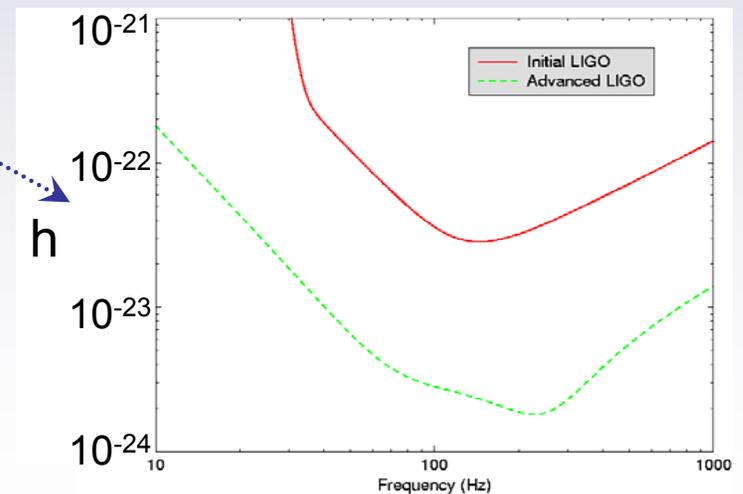
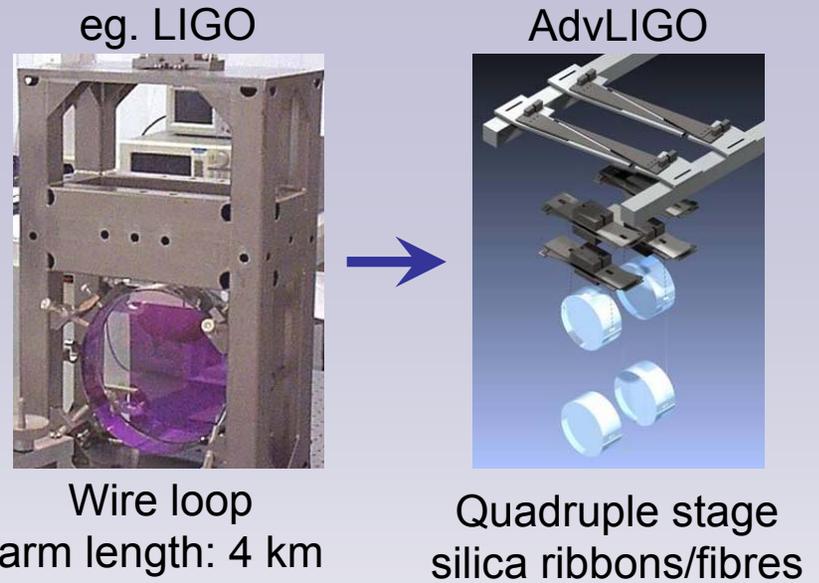
Silica fibres are welded to the ears in the completion of the lower-stage of the GEO600 mirror suspension.



# Investigations into hydroxy-catalysis bonding

## Planned applications

- The planned upgrades for AdvLIGO and Advanced VIRGO plan to incorporate the GEO600 technology for significantly improved thermal noise performance (in addition to other improvements, e.g higher power lasers).
- Construction of the ultra-rigid, ultra-stable optical benches for LISA.



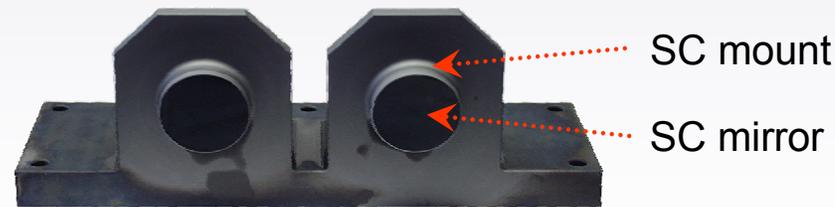
Design sensitivity curves for the LIGO and AdvLIGO detectors.



# Catalysis bonding settling time dependence

## Motivation

- Many optical systems layouts have stringent requirements for strength, rigidity, stability and **alignment**.
- Hydroxy-catalysis bonding fulfills all these requirement.
- One possible disadvantage of this technique is that the time taken for a typical bond to “set” at room temperature is in the region of a few tens of seconds. This only allows a **short period of time in which to align** the various components on the optical bench
  - Glasgow has investigated how to extend the settling time of hydroxy-catalysis bonds through varying the hydroxide concentration and lowering the temperature.  
(Reid et al, Physics Letter A, 2007)



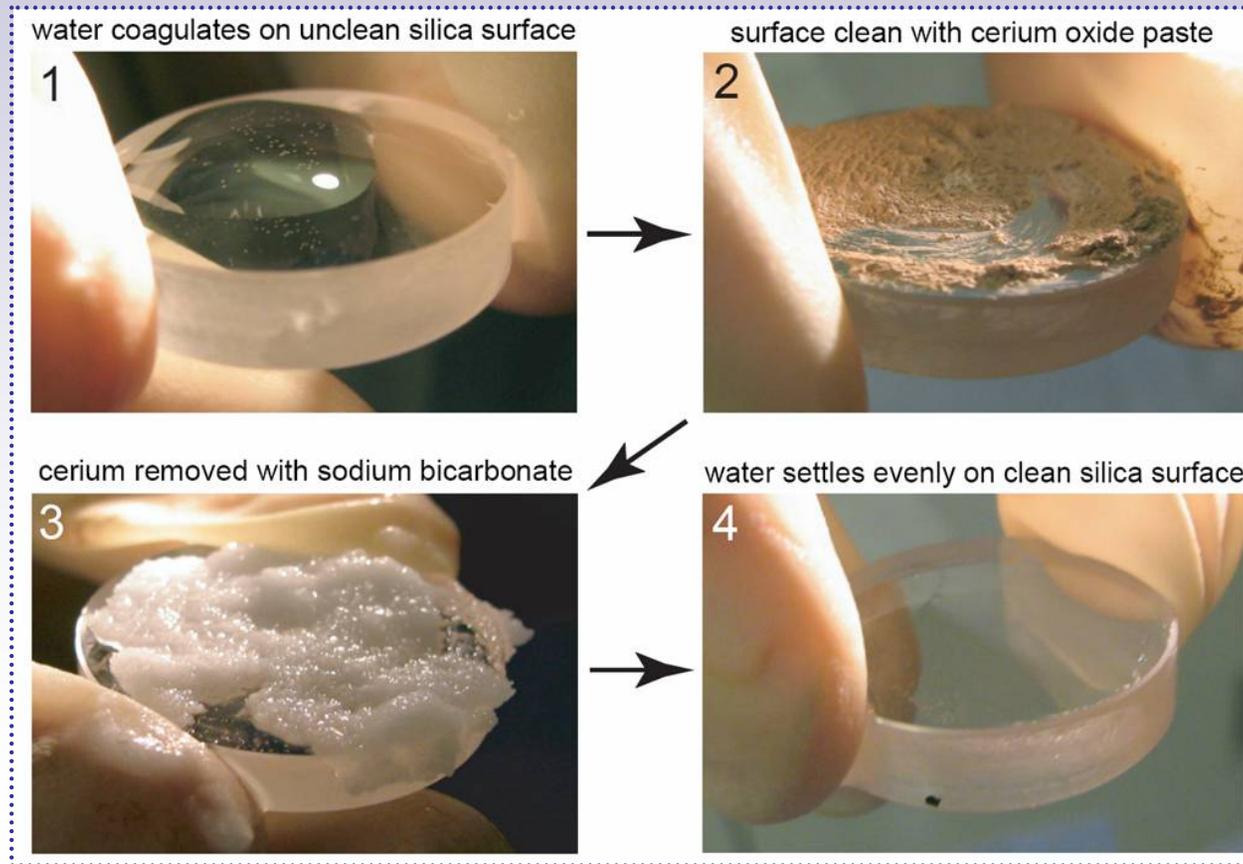
Picture of silicon carbide mirrors bonded to a silicon carbide base mount for GAIA



# The processes involved in hydroxy-catalysis bonding

## Surface preparation

- Contaminants on the silica surface will likely inhibit hydration. The silica surfaces to be jointed are thus taken through a cleaning process to remove any contaminants and to ensure maximum hydration:

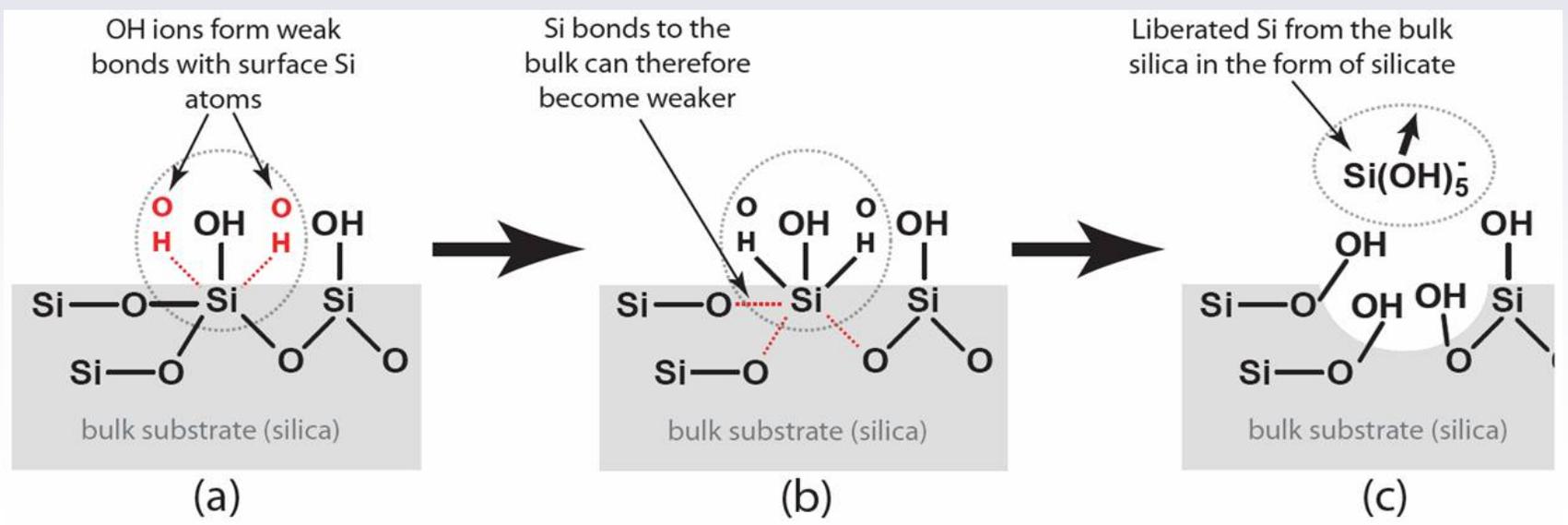




# The processes involved in hydroxy-catalysis bonding

## Etching — Etching of the silica surfaces to be bonded

- Placing a solution with a high concentration of OH<sup>-</sup> ions on the surface of silica causes etching to take place.
- Free OH<sup>-</sup> ions form weak bonds with silicon atoms on the substrate surface causing the original lattice **bonds to weaken**
- It becomes possible for the silicate molecule to break away from the bulk structure, producing **Si(OH)<sub>5</sub><sup>-</sup> molecules in solution**.

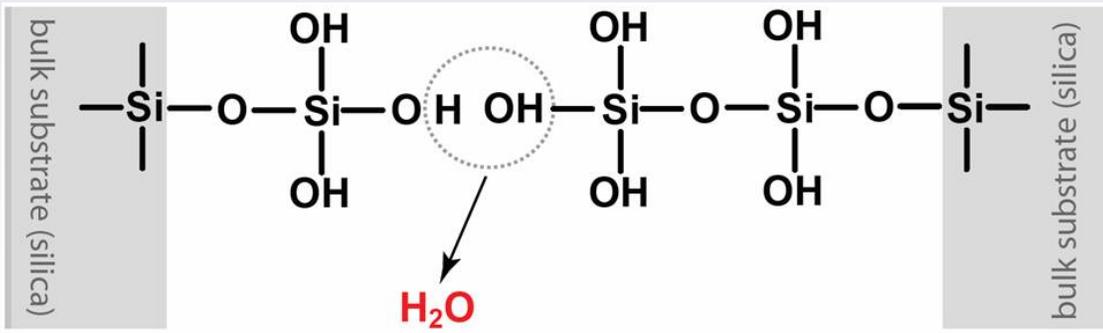
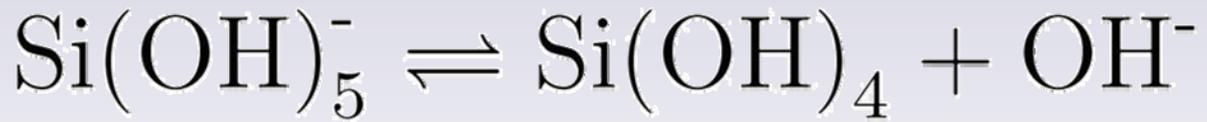




# The processes involved in hydroxy-catalysis bonding

## Bonding – Polymerisation of silicate in solution

- However, below pH 11, the silicate ion hydrolyses to soluble  $\text{Si(OH)}_4$  and  $\text{OH}^-$ . When the concentration of  $\text{Si(OH)}_4$  molecules reaches 1→2%, the solution polymerises and becomes “rigid” (R.K. Iler, 1979, *The Chemistry of Silica*).
- $\text{Si(OH)}_4$  is a **monomer** which likes to form a **polymer** arrangement:



- Bonding starts to dry and evaporates  $\text{H}_2\text{O}$
- Bond thickness ~ 100 nm





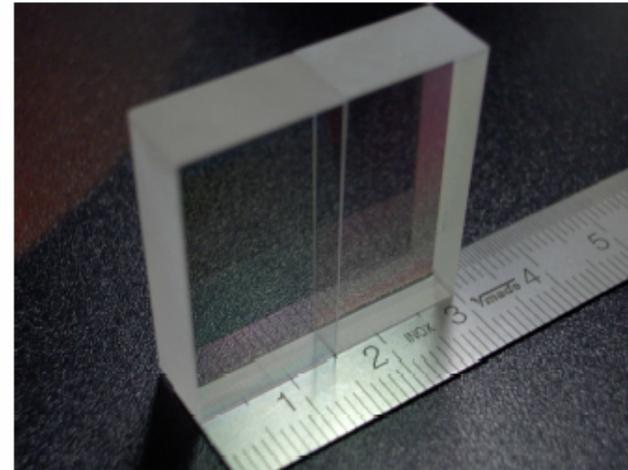
# Investigations into hydroxy-catalysis bonding

Development  
Theory of silicate bonding  
Cleaning process for bonding

Material for silicate bonding  
Clean room

## De-Bonding

- De-bonding is possible within a few hours from bonding, depending on size, shape and quality of the bond
- Using ultra-sonic bath with detergent solution (10% Decon<sup>®</sup>)
- Each bonding and de-bonding process can slightly damage the surface for bonding





# Investigations into hydroxy-catalysis bonding

## Ongoing work in Glasgow includes:

- Further investigation of bonding samples with a **ground finish**.
  - align optics **without danger of optical contacting**
  - **no time constraint** on achieving alignment
  - apply bonding solution with pieces in situ

- Bonding other materials

e.g. Silicon-silicon bonded samples  
for verifying the feasibility of monolithic  
silicon suspensions for 3<sup>rd</sup> generation  
gravitational wave detectors:

- **mechanical strength**
- **thermal conductivity (at low T)**
- **mechanical loss**



Silicon-silica hydroxy-catalysis bond



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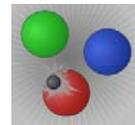
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UK Astronomy Technology Centre, Edinburgh



Wed Feb 21st 2007



Dr. Calum I. Torrie

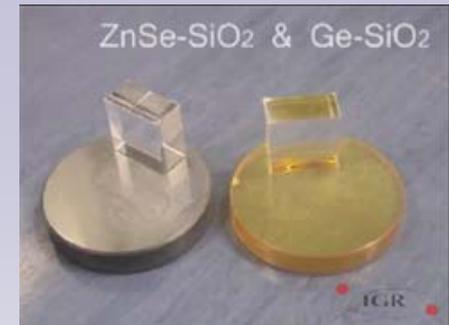


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# Investigations into hydroxy-catalysis bonding

- Successful test for Astrium D for feasibility of ZnSe/SiO<sub>2</sub> and Ge/SiO<sub>2</sub> bonds

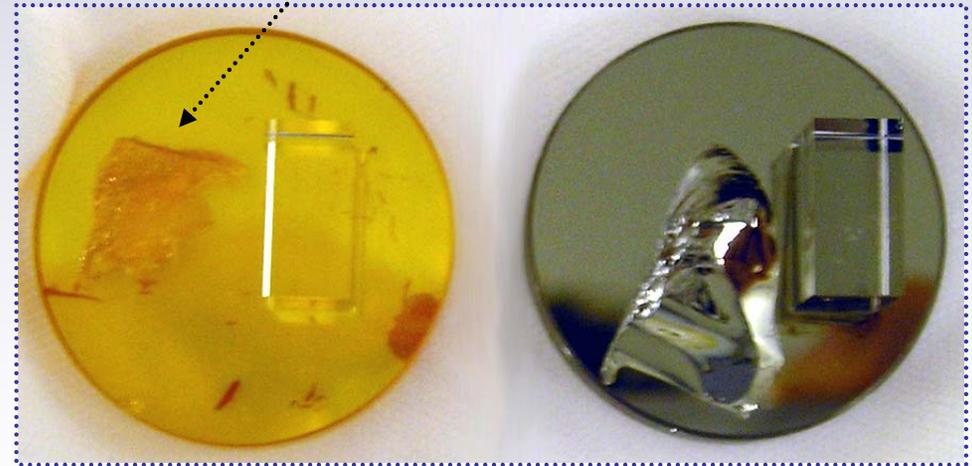


Temperature cycling:  
SiO<sub>2</sub> (silica) bonded to  
ZnSe (zinc selenide)  
and Ge (germanium)  
down to 77K



Images from video of temperature cycling of SiO<sub>2</sub> bonded to ZnSe down to 77K

Damage due mechanical strength testing.



SiO<sub>2</sub> bonded to ZnSe

SiO<sub>2</sub> bonded to Ge



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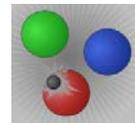
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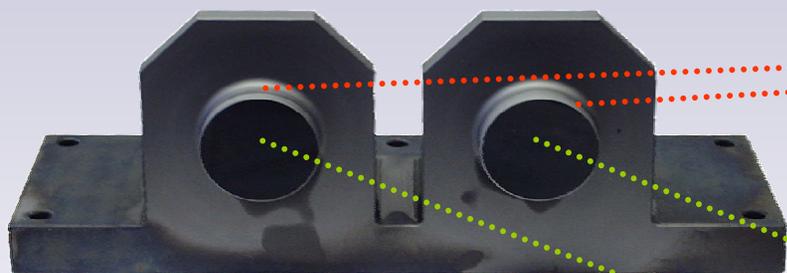


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# Investigations into hydroxy-catalysis bonding

- Astrium France – IGR subcontract – silicon carbide optical assemblies for GAIA – **patent application filed**
- **Current non-disclosure agreement in place** with TNO-TPD for further evaluation

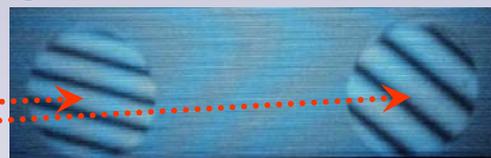


Picture of silicon carbide mirrors bonded to a silicon carbide base mount for GAIA

**No relative tilt change** between number of fringes within the resolution of the interferometer

Relative tilt in mirrors before and after bonding **< 1 arcsec** (well within design specification of **3 arcsecs** in relative tilt).

fringe pattern of SiC optical mount

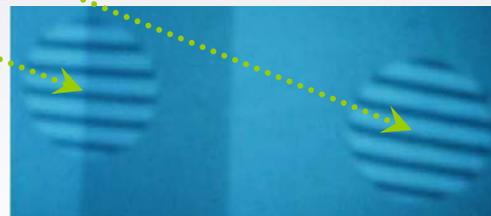


before oxidation



after oxidation

fringe pattern of mounted SiC mirrors



mirrors mounted dry



mirrors hydroxy-catalysis bonded



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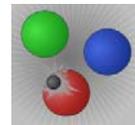
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# TEOPS Activity

- Submitted (Proposals)
  - European Gravitational Wave Observatory (EGO)
    - silica ribbons for monolithic suspensions - submitted
  - TEOPS Proposal part of larger PPARC proposal for ELT
    - Development of key passive & adaptive mirror technologies using SiC mirrors
  - ATLAS upgrade
    - Development of supermodule structures
- Discussions (Contract)
  - Advanced LIGO
    - Optical and Mechanical Systems
  - GEO 600 and Hannover AEI Prototype
    - Upgrades to existing detector
    - Isolation and Suspension design
- Other
  - Adam Smith Scholarships
    - Mech Eng & Chemistry
  - PIPSS & mini-PIPSS
  - British Consulate LA and Scottish Executive
    - Caltech / Glasgow (SUPA) expansion

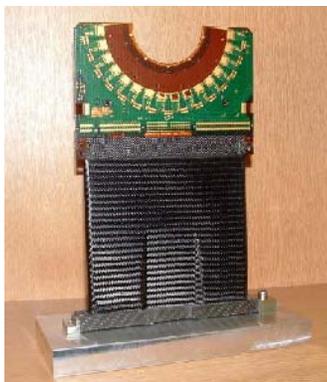
# Commonality

- Significant areas of commonality: some examples:
  - Cryogenics
    - The ATC has decades of experience in constructing reliable and robust instruments operating at cryogenic temperatures (as low as 4 K and even below 100 mK)
  - Cryogenic operation is now of interest for future generations of both gravitational wave detectors and colliders for particle physics



# Commonality

- New materials
  - e.g. silicon-carbide being looked for use in astronomical instruments, particle physics detectors and in gravitational wave detectors
  - CSiC trials underway in Glasgow
    - Thermal, mechanical, vacuum & bonding



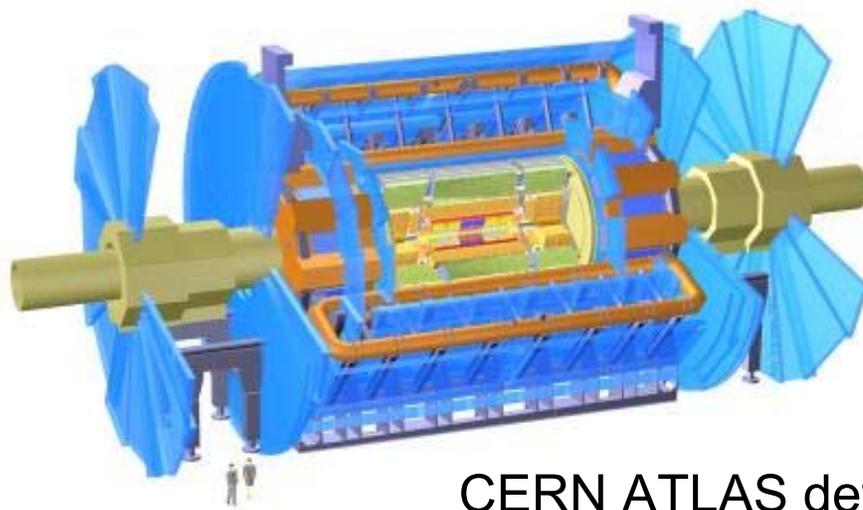
LHC module using carbon fibre mount  
– Si-C considered as replacement



Si-C lightweighted telescope mirror  
(courtesy M. Krodel)

# Commonality

- Particle physics detector groups have experience in constructing detectors on an “industrial” scale
  - will be required in astronomy as telescopes increase in size (and number of telescopes in the case of arrays)



CERN ATLAS detector